





# The VGP News



"Mind and Matter and Crucible"  
(Originally appeared in *The Geochronology*, 9, 5, 1985.)

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## News & Announcements

### Travel Grants to IGC

The U.S. National Committee for Geochemistry is seeking funding for its Travel Grant Program to the 27th International Geological Congress to be held in Moscow August 4-14, 1984. In cooperation with other organizations, the committee seeks to ensure appropriate U.S. participation by providing 10-20 travel grants to enable geochemists residing in the United States to attend. Travel grants are to be awarded based in part on a screening subcommittee's ranking of abstracts submitted for presentation at the congress. Special consideration will be given to younger geochemists and those judged to benefit most by participating in this important international congress.

To apply, send six fastened-together sets of completed applications, including completed form and abstract (500-1200 words) of paper to be presented, to W. L. Petrie, USNG/Geochemistry, NAS-NRC, 2101 Constitution Ave., Washington, DC 20418 by January 31, 1984. (Firms available from Petrie.) Travel grants must use U.S. flag carriers wherever possible and most like a meaningful trip report before October 14, 1984.

Depending on the availability of funds, travel grant awards may be made by May 1, 1984. However, cancellations and other factors may delay a few grant awards to as late as August 1, 1984, or possibly after the congress.

### Research Grants Announced

Kraeger Enterprises, Inc., has announced the winners of its 1983 Geochron Research Competition. Two awards were granted to support research proposals in each of the following areas: K-Ar dating, C-14 analyses, and Stable Isotope Ratio Analyses (SIRA).

Winners in the K-Ar dating area are James J. Hardy, Jr., Northern Arizona Univ., The use of the K-Ar method to date a major thrust event in west-central Arizona; and Christopher S. Lynnes, Univ. of Michigan, Correlating

tion with age of magnetization in Cambro-Ordovician intrusives from Colorado.

In C-14 analyses: Kee Hynn Kim, Florida State Univ., Cross-check of uranium-series disequilibrium ages by radiocarbon dating; Marine phosphate nodules and their associated sediments; and Charles K. Paul, Scripps Institution of Oceanography. The origin of stratigraphic offsets in deep-sea cores.

In SIRA: Teofilo A. Abramo, Jr., Washington Univ., Origin and significance of sulfide phases in selected mantle assemblages; and Virginia B. Sisson, Princeton Univ., Oxygen isotope work on the Ponder pluton and associated rocks.

According to Krueger, the awards, which amount to about \$1500 worth of analytical services for use in the research, are to be made again in 1984 in the same three areas as well as in Rb-Sr or U-Pb analyses. More information may be obtained from Krueger Enterprises, Inc., 24 Blackstone St., Cambridge, MA 02139 (telephone 617-876-3691).

## Meetings

### Cosmogenic Radionuclides

Cosmic rays interact with the earth's atmosphere and surface to produce the "cosmogenic" nuclides. In many instances the radioactive ones are readily distinguished from the anthropogenic and meteoric backgrounds. Measurements of these cosmogenic radionuclides (RCN) can contribute to the solution of a variety of geophysical problems (Lal and Peters, 1967). Recent progress in this area was discussed at a symposium entitled "Application of Cosmic-Ray-Produced Nuclides in Geophysics" held May 30, 1983, at the AGU Spring Meeting in Baltimore (see *Eos*, May 3, 1983, pp. 289-294, for the abstracts). We summarize here the symposium presentations.

The RCN accumulate differently in different terrestrial reservoirs. Table 1 shows some estimated global production rates and abundance levels for selected samples. In relatively simple collectors such as ice, the measurements may shed light on variations in production rates. In other collectors such as manganese nodules the observations may tell more about the object than about cosmic rays. Accordingly, we divide the following text into two sections. The first summarizes symposium contributions that emphasized cosmic ray histories and the second those that focused on the sample. Where the distinction blurs, the two sections overlap.

### Cosmic Ray Variations

The total flux of solar and galactic cosmic ray particles determines the production rates of the RCN. J. R. Jokipii reviewed the types of temporal variability exhibited by the cosmic-ray flux. He noted the absence of large (2-3-fold), long-term ( $10^3$ - $10^5$ -years) variations that might be associated with galactic processes. Changes in solar activity are known to induce shorter-term variations. Of special interest here are (1) the 11-year solar cycle which normally gives rise to 25% changes in the cosmic ray flux and (2) a "cycle" lasting perhaps 200-400 years during which the flux may increase by a factor of 2 or 3 in response to solar modulation. Superimposed on the above is a cycle of about 10<sup>5</sup> years attributed to changes in the geomagnetic field (Björck, 1967). Direct observation of the sun has established the 11-year solar cycle and shorter

TABLE 1. Cosmogenic Radionuclides in Selected Samples

Isotope	$t_{1/2}$ , years	Production rate, atom $\text{cm}^{-2} \text{s}^{-1}$	Sample	Concentration, atom $\text{g}^{-1}$
$^{10}\text{Be}$	$2.7 \times 10^6$	$\sim 5 \times 10^{-7}$	Seawater	6
$^{14}\text{C}$	$5.7 \times 10^3$	$1.57 \times 10^{-12}$	Corals	$7 \times 10^6$
$^{36}\text{Cl}$	$3.0 \times 10^5$	$1.1 \times 10^{-13}$	Groundwater	$10^4$
			Ice	$4 \times 10^5$
			Tektites	$2 \times 10^6$
			Rain	$1.6 \times 10^4$
			Soils	$10^3$ - $10^4$
			Lavas	$5 \times 10^4$
			Phosphorites	$10^5$
			Aln nodules	$5 \times 10^4$
			Petroleum	$5 \times 10^4$

\*Suess [1980].

\*\*Reiss et al. [1981].

term variations, namely bubbling and gusting of the solar wind. Forbush decreases, and solar flares. Measurements of the RCN may supplement these measurements. They take on primary importance as we go further back in time.

### $^{14}\text{C}$ in Corals

E. Druffel summarized some studies of  $^{14}\text{C}$  in corals. The 11-year solar cycle leads to calculated variations in the  $^{14}\text{C}/^{12}\text{C}$  ratio too small to detect with available methods. Variations in  $^{14}\text{C}/^{12}\text{C}$  ratios with a period of 200-400 years have been documented in several laboratories. Druffel has shown that corals that grew during the "little ice age" have elevated  $^{14}\text{C}$  contents.

There are three questions now under study. (1) What is the relationship between climate and solar activity as recorded by RCN's? Present evidence is contradictory. (2) To what extent do oceanic circulation patterns influence the  $^{14}\text{C}/^{12}\text{C}$  ratio of the atmosphere? Evidence from corals suggests that the effect is small. (3) Has  $^{14}\text{C}$  production and its exchange rate between air and sea changed much over the last 10,000 years? At present, unknown factors cause  $^{14}\text{C}$  ages to be 10% younger than U-series ages of Pleistocene samples. The difference could reflect either an increased  $^{14}\text{C}$  production rate or a lower partial pressure of  $\text{CO}_2$  in the atmosphere.

### RCN in Ice

K. Nishizumi determinations of longer-lived cosmogenic radionuclides in ice. Unlike soils or sediments, ice cores provide samples minimally diluted by stable isotopes that may interfere with measurements. In favorable cases ice cores give sharp temporal resolution, too. On the other hand, the RCN deposition rates can vary in response to atmospheric processes that may be only poorly understood. One way to test the fidelity of an isotope record in ice is to compare it to one in a different collector, such as  $^{14}\text{C}$  in tree rings.

Measurements of both  $^{14}\text{C}$  and  $^{10}\text{Be}$  in the Dye 3 ice core correlate well with sunspot activity subsequent to 1950. The  $^{10}\text{Be}$  contents change less because the isotope's longer atmospheric residence time dampens oscillations due to production rate variations. For the period prior to 1950 agreement between the  $^{14}\text{C}$  and  $^{10}\text{Be}$  is poor, possibly due to unknown reasons. Raisbeck et al. (1981) analyzed  $^{10}\text{Be}$  in the Dome C ice core and found that samples deposited during the Matuyama minimum had elevated  $^{10}\text{Be}$  contents as expected. Raisbeck et al. (1981) and Oeschger and coworkers have reported increased  $^{10}\text{Be}$  concentration during the last ice age. The  $^{10}\text{Be}$  profiles correlate well with  $\text{CO}_2$  results. It is not known whether the enhanced  $^{10}\text{Be}$  deposition reflects lower precipitation rates, changes in atmospheric circulation patterns or a higher production rate.

Finally, Nishizumi summarized recent progress in the  $^{10}\text{Be}/^{14}\text{C}$  dating of ice samples. Relative ages for two samples A and B can be calculated by assuming a constant  $^{10}\text{Be}/^{14}\text{C}$  ratio in precipitation and the absence of radionuclide transport in ice. With the adoption of a modern  $^{10}\text{Be}/^{14}\text{C}$  atom-atom ratio of 8, Nishizumi concludes that near-surface ice from the Allan Hills region of Antarctica has an age of about  $5 \times 10^5$  years.

### $^{10}\text{Be}$ in Lake Sediments

M. Wahlen reported the  $^{10}\text{Be}$  contents of varved sediments from two New York State lakes with minimal bioturbation. The purpose of the study was to search for long-term short-term variations in the rates of production. The  $^{10}\text{Be}$  fluxes calculated for Green Lake are divided into three periods: a baseline period with low deposition rates ( $0.016$  atom  $\text{cm}^{-2} \text{s}^{-1}$ ), a period of enhanced rates during the Spörer and Maunder minima (up to  $0.1$  atom  $\text{cm}^{-2} \text{s}^{-1}$ ), and the last 200 years or so during which  $^{10}\text{Be}$  fluxes have steadily climbed toward a value near  $0.3$  atom  $\text{cm}^{-2} \text{s}^{-1}$ . The observations for the second period match in size but not in temporal detail with

lar results from  $^{14}\text{C}$ . The recent increase in  $^{10}\text{Be}$  deposition is attributed to increased soil erosion.

### Sample Oriented Studies

#### $^{10}\text{Be}$ in Rainfall

Deposition rates for the longer lived RCN are not yet well known. O'Brien [1979] calculated a global average production rate of  $0.025$  atom  $\text{cm}^{-2} \text{s}^{-1}$  for  $^{10}\text{Be}$ , while Reiss et al. (1981) estimated  $0.021$ .

C. J. Stensland and M. C. Monaghan independently presented the results of several determinations of  $^{10}\text{Be}$  in rainwater. The authors designed the experiments to learn more about the factors influencing  $^{10}\text{Be}$  deposition. Stensland's group deployed traps open only when rain fell. At one site (Boulder, Ill.) the average monthly fall-out rate ranged from about  $1 \times 10^{-4}$  to  $2 \times 10^{-4}$  atom  $\text{Be} \text{ g}^{-1} \text{ H}_2\text{O}$ . Values were highest from May to August and lowest from October through March. The variations parallel those observed for  $^{10}\text{Be}$ . Stensland noted that the Ca contents of the rainwater samples indicate contamination by true soil particles. Raindrops or wind presumably lifted these particles into the air where they were entrained by raindrops on their way to the collectors. The concentration of  $^{10}\text{Be}$  in the soil particles was not measured directly. Stensland estimates the contribution from this source may be as high as  $1 \times 10^4$  at  $^{10}\text{Be} \text{ g}^{-1} \text{ H}_2\text{O}$ . This work demonstrates the need for care in the collection of samples.

Monaghan attempted to estimate the annual global production of  $^{10}\text{Be}$  from measurements of  $^{10}\text{Be}$  and  $^{36}\text{Cl}$  in rainwater. Each sample consisted of a year's precipitation collected at one of eight sites in the United States. Dust may have contaminated some samples. The observed, latitude-dependent fluxes of  $^{10}\text{Be}$  were normalized to global ones with the aid of the  $^{36}\text{Cl}$  measurements and the known global deposition rate of  $^{36}\text{Cl}$ . With various assumptions and a theoretical correction for tropospheric production Monaghan obtains a global production rate of about  $0.018$  at  $^{10}\text{Be} \text{ cm}^{-2} \text{ s}^{-1}$ . From a comparison of the different sites, Monaghan concludes that the  $^{10}\text{Be}$  flux as measured by precipitation at a particular location may not reliably reflect the  $^{10}\text{Be}$  production rate.

### $^{36}\text{Cl}$ in Seawater

R. D. Willis presented the first profile of  $^{36}\text{Cl}$  in seawater. The  $^{36}\text{Cl}$  half-life (289 years) makes this isotope important for treatments of oceanic circulation. The time scale for ventilation of the ocean is now estimated from determinations of the less sensitive  $^{10}\text{Be}$ . The samples discussed by Willis were collected at depths between 0 and 600 m in the north Pacific (GEOSECS II). Each of the five samples weighed about 2 tonnes (t), was five samples weighed about 2 tonnes (t), was counted for 2 months and contained on the order of only  $6 \times 10^4$  atom  $\text{g}^{-1} \text{ H}_2\text{O}$ . The  $^{36}\text{Cl}$  data reported for 600 m differs significantly from the prediction of a  $^{36}\text{Cl}$  box diffusion model. Three times lower suggests a removal rate of  $^{36}\text{Cl}$  some 8 times faster than expected. Samples from depths of 4000 m remain to be counted.

### Sources of Atmospheric Methane

S. Hameed estimated from the carbon isotopic ratios measured in atmospheric methane the fraction of methane derived from (1) the incomplete combustion of fossil fuels, (2) the partial burning of biomass, and (3) bacterial sources. Biomass burning may be more important than previously realized.

### $^{10}\text{Be}$ in Lavas

L. Brown discussed progress in studying  $^{10}\text{Be}$  in volcanic rocks. The primary purpose of this ongoing work is to demonstrate that island-arc lavas contain measurable traces of  $^{10}\text{Be}$ . Pelagic sediments subducted from the trench, Semprum lavas, estimates of the fraction of sediment-derived material present in a sample may then be possible.  $^{10}\text{Be}$  was a good tracer for sedimentary matter for the following reason: Ocean sediments may con-

tain up to  $10^{10}$  atom  $^{10}\text{Be} \text{ g}^{-1}$  while igneous rocks contain a few million years old  $^{10}\text{Be}$  should have virtually none. In the last year Brown and his colleagues have doubled the number of analyses of volcanic rocks, bringing the total to about 40. The experiments took pains to try to minimize contamination from soil and rain. They divide their results into three groups. Twenty of 21 samples from the Aleutians and Central America contain between  $2 \times 10^4$  and  $5 \times 10^6$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . In contrast, 10 non-island arc lavas and  $^{10}\text{Be}$  island arc samples from the Andes and the Far East seem to constitute an intermediate group. Most have low  $^{10}\text{Be}$  contents but a few are similar to the Central American lavas. Brown presents evidence that the  $^{10}\text{Be}$  systematics are more firmly established. For the present he concludes that island-arc lavas do indeed contain  $^{10}\text{Be}$  that once resided in pelagic ooze.

### $^{10}\text{Be}$ in Phosphorites, Aln Nodules, and Petroleum

C. Tuzi reported measurements of  $^{10}\text{Be}$  in marine phosphorites. The aim of the study was to learn more about the timing of phosphorite formation. W. C. Burnett and H. H. Yeh have argued in several publications that  $^{10}\text{Be}$  disequilibrium dating supports a recent ( $2 \times 10^5$  years) origin for certain phosphorites collected off the coasts of Chile and Peru. Others have questioned this conclusion. Tuzi and coworkers showed that 11 phosphorites contain about  $2 \times 10^4$  atom  $^{10}\text{Be} \text{ g}^{-1}$ , about twice as much as nearby sediments and  $^{21}\text{Bi}$ -dated phosphorites with  $^{234}\text{U}/^{238}\text{U}$  top-to-bottom age differences of a few thousand years have no  $^{10}\text{Be}$  variability within experimental error as would be expected if the  $^{10}\text{Be}$  ages were valid. Two kinds of tanks were discussed. Five million-year-old phosphorites from Bone Valley, Fla., have low  $^{10}\text{Be}$  contents, consistent with radiocarbon dating, or a low initial complement of  $^{10}\text{Be}$  remains to be seen whether very old submarine samples contain  $^{10}\text{Be}$  from water-sediment infiltration. A negative finding would strengthen the case for recent phosphorite origin.

Several groups have exploited accelerator mass spectrometry for the measurement of  $^{10}\text{Be}$  and other cosmogenic radionuclides in manganese nodules and crusts. The results may be used to calculate growth rates. They may also give information about the conditions under which deposition occurred. A. Almagi and coworkers studied a manganese crust some 30-40 cm thick dredged from a depth of 4830 m in the central Pacific. They measured the  $^{10}\text{Be}$  and metal contents of 18 samples. A plot of  $^{10}\text{Be}$  by depth suggests a two-stage accumulation history: an early one lasting from 11 to 6 m.y. B.P. with a deposition rate of  $4.8 \text{ mm/MY}$  and a late one with a shift at 6 m.y. B.P. may coincide with a discontinuity in the  $^{10}\text{Be}$  record noted by others. With the aid of the growth rates, Almagi calculated dates for several petrographic and chemical changes observed in the crust. Many of these events apparently coincide with changes in paleoclimatic circulation inferred independently from the study of deep-sea sediments. Almagi concludes that the compact Mn nodules and crusts can provide paleoclimatic histories.

Motivated by the lack of reliable isotopic techniques for the age-determination of petroleum, F. Yon and coworkers have begun a survey of  $^{10}\text{Be}$  in various petroleum reservoirs. Oils older than 25 m.y. contain undetectable  $^{10}\text{Be}$ ; crude oils thought to be 2-5 m.y. old have detectable quantities of  $^{10}\text{Be}$ ,  $4 \times 10^4$  atom  $\text{g}^{-1}$ ; a modern, hydrocarbon-rich sediment contains  $4 \times 10^4$  atom  $^{10}\text{Be} \text{ g}^{-1}$ , some of which may attach to the elastic component. The authors note that many factors may influence the  $^{10}\text{Be}$  contents of crude oils, among them initial deposition conditions, exchange with water, sediments, reservoir rocks and kerogen, and migration history.

### $^{10}\text{Be}$ in Soils

M. Pavich and J. Klein both reported on  $^{10}\text{Be}$  measurements in soils. The multiple origins of soil particles pose a difficult problem for isotopic dating of samples more than  $10^3$  years old. It has been suggested that with suitable modeling,  $^{10}\text{Be}$  measurements may provide useful information about soil chronology. Pavich measured the  $^{10}\text{Be}$  contents in soil columns from Merced River, Calif., terraces ranging in age from 0.04 to 3 m.y. As expected, total  $^{10}\text{Be}$  inventories increase with column age and thickness. For the oldest columns, however, the standing crop of  $^{10}\text{Be}$  falls short of the amount anticipated based on current estimates of the deposition rate and the assumption of complete retention. From the shortfall the authors estimate a  $^{10}\text{Be}$  residence time of  $8 \times 10^5$  years. Monaghan et al. (1983) obtain a lower value of about  $10^6$  years from an analogous study of other soil columns. Whether the discrepancy arises from sampling artifacts or site-specific differences is not resolved. Pavich uses a correlation between clay and  $^{10}\text{Be}$  contents in the soils.

Klein and coworkers have been studying erosion in various watershed regions by applying mass balance to  $^{10}\text{Be}$ . They measure the  $^{10}\text{Be}$  contents of sediments from rivers or on continental margins, multiply by the sediment flux, and compare the result with the fall-out rate of  $^{10}\text{Be}$  over the watershed. In a few cases such as the Amazon and the Susquehanna rivers,  $^{10}\text{Be}$  inflow and outflow match to within a factor of 2. In many others such as the Mississippi and Yangtze rivers,  $^{10}\text{Be}$  loss may exceed accretion by as much as a factor of 10. No "primordial" watershed has been found to give a steady-state  $^{10}\text{Be}$ .

Thus, increased erosion due to agriculture may account for part but need not explain all of the net loss. Klein notes that, with a few exceptions, areas with the highest erosion rates produced sediments with the highest  $^{10}\text{Be}$  contents. The authors conclude that erosion tends preferentially to carry away particles rich in  $^{10}\text{Be}$ .

### $^{36}\text{Cl}$ in Groundwater

The relative inertness of aqueous  $\text{Cl}^-$  renders it especially suitable as a hydrologic tracer. Groundwater contains  $^{36}\text{Cl}$  derived from three main sources: cosmic ray bombardment of the air and of rocks, activation of rocks by fission-produced neutrons, and bomb testing. With appropriate modeling the  $^{36}\text{Cl}$  contents of groundwater systems give chronological and hydrodynamic information. Accelerator mass spectrometry makes it possible to measure the low levels of  $^{36}\text{Cl}$  encountered. S. N. Davis, H. W. Bendy, and P. L. Airey discussed the application of  $^{36}\text{Cl}$  measurements to the study of the Great Artesian Basin in Australia and to the Milk River Aquifer in Alberta, Canada.

In the Great Artesian Basin, the decrease systematically from the source to the discharge region. With minor exceptions, the isochrons mapped from the  $^{36}\text{Cl}$  ages agree well with isochrons independently calculated from the known hydraulic parameters of the system. The authors interpret the agreement as a strong endorsement of the assumptions of  $^{36}\text{Cl}$  groundwater dating.

The Milk River Aquifer behaves differently:  $\text{Cl}^-$  and  $^{36}\text{Cl}$  contents increase with distance from the recharge or source region. While the  $^{36}\text{Cl}/^{35}\text{Cl}$  ratio decreases, the  $^{36}\text{Cl}$  contents increase. The concentration increases in terms of "ion filtration." The  $^{36}\text{Cl}/^{35}\text{Cl}$  ratios from yield water ages that exceed estimates from numerical modeling. The flow pattern for glacial interruption of the flow pattern for long periods of time. Some alluvial  $^{36}\text{Cl}$  by contemporary dissolution of ancient chloride de-

positis may also occur. This process, however, would increase the calculated  $^{36}\text{Cl}$  ages.

### $^{10}\text{Be}$ & $^{26}\text{Al}$ in Impact Ejecta

The application of cosmogenic radionuclides to the study of impact-produced materials was discussed by G. Raisbeck and F. Tera. Last year Pal et al. [1982] showed that australasian tektites contain  $\sim 2 \times 10^4$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . They argued that the  $^{10}\text{Be}$  and the tektites formed on earth, but could not entirely exclude the possibility that a small extraterrestrial component, either lunar or meteoritic, carried some  $^{10}\text{Be}$ . If the  $^{10}\text{Be}$  were extraterrestrial, however, the tektites would also have to contain measurable quantities of other cosmogenic radionuclides such as  $^{26}\text{Al}$  and  $^{26}\text{Mn}$ . Raisbeck used accelerator mass spectrometry to measure the  $^{26}\text{Al}$  contents of several australasian tektites and other impact-related objects. The  $^{26}\text{Al}/^{10}\text{Be}$  ratios observed were consistent with a terrestrial but not an extraterrestrial origin. These observations unambiguously rule out a lunar origin for tektites. Raisbeck also found that other impact-produced materials such as Libyan desert glass contain  $^{26}\text{Al}$  in amounts consistent with what cosmic rays would produce on bombarding terrestrial surface matter for  $\sim 10^7$  years. He suggested, as have others, that the cosmogenic radionuclides may prove useful in the calculation of surface exposure ages of terrestrial rocks.

Tera presented the  $^{10}\text{Be}$  contents of more than 80 australasian tektites and the  $^{26}\text{Al}$  contents of six. Again the  $^{26}\text{Al}/^{10}\text{Be}$  ratios were terrestrial. Tera assessed the small range of values observed for  $^{10}\text{Be}$ , values which average about  $2 \times 10^4$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . Sediments harvested from continental margins exhibit a similar, small range of  $^{10}\text{Be}$  contents. Soils, in contrast, have  $^{10}\text{Be}$  contents that differ markedly from location to location, ranging from  $1 \times 10^4$  to over  $10^6$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . Tera suggests a continental margin for the site of the impact that produced australasian tektites. An origin on a continental margin would accommodate the difficulty specialists have had in identifying a source crater: the crater would have been covered. A margin site may also furnish material from considerable depth with appreciable  $^{10}\text{Be}$  content and thus would help explain the mid-ranges of the data.

### Prospects

The advent of accelerator mass spectrometry has greatly facilitated the measurement of such long-lived beta-emitters as  $^{36}\text{Cl}$ ,  $^{26}\text{Al}$ ,  $^{10}\text{Be}$ , and  $^{14}\text{C}$ . The new technique obviates the need for tedious radiochemistry, reduces by 1000 times the minimum sample masses, and gives results in hours rather than days or weeks of counting. Technical advances continue. D. Elmore pointed to prospects for measuring  $^{10}\text{Be}$  and certain stable elements; for example, platinum. Accelerator mass spectrometry may prove helpful in the measurement of solar-neutrino-produced isotopes and in other problems in particle physics that have geologic detectors. Willis outlined progress made by Hurst and coworkers at Oak Ridge in the detection of  $^{81}\text{Kr}$  by laser-assisted separation and mass spectrometry. The advance of resonance ionization techniques promises a host of new applications.

K. Turekian sounded several cautionary notes. He pointed out that because local deposition rates of the RCN may vary appreciably, certain investigations will require large numbers of measurements to establish clear patterns in the results. The high cost of using the new measurement techniques militates against such surveys except where a cheaper approach exists. Furthermore, gaps in our understanding of the geochemistry of the RCN and of the systems in which they are studied may cloud the interpretation of even the most precise results. For example, diffusion as well as successive lunar deposition may influence the  $^{10}\text{Be}$  contents of Mn crusts and nodules.

The foregoing text indicates the wide variety of geophysical problems related to cosmogenic radionuclides. While the less well-known isotopes  $^{10}\text{Be}$ ,  $^{26}\text{Al}$ ,  $^{36}\text{Cl}$ , and  $^{14}\text{C}$  seem unlikely to challenge  $^{14}\text{C}$  in importance, they occupy niches that  $^{14}\text{C}$  cannot fill. We regard the study of  $^{10}\text{Be}$  in volcanic rock as fundamental in importance. The hydrogeological measurements of  $^{36}\text{Cl}$  have begun to meet expectations. The analysis of cosmogenic radionuclides in various terrestrial reservoirs augments our knowledge of sedimentary, depositional, and erosional processes. Much of the work described is exploratory. We anticipate that rapid progress in the area will continue.

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### $^{10}\text{Be}$ & $^{26}\text{Al}$ in Impact Ejecta

The application of cosmogenic radionuclides to the study of impact-produced materials was discussed by G. Raisbeck and F. Tera. Last year Pal et al. [1982] showed that australasian tektites contain  $\sim 2 \times 10^4$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . They argued that the  $^{10}\text{Be}$  and the tektites formed on earth, but could not entirely exclude the possibility that a small extraterrestrial component, either lunar or meteoritic, carried some  $^{10}\text{Be}$ . If the  $^{10}\text{Be}$  were extraterrestrial, however, the tektites would also have to contain measurable quantities of other cosmogenic radionuclides such as  $^{26}\text{Al}$  and  $^{26}\text{Mn}$ . Raisbeck used accelerator mass spectrometry to measure the  $^{26}\text{Al}$  contents of several australasian tektites and other impact-related objects. The  $^{26}\text{Al}/^{10}\text{Be}$  ratios observed were consistent with a terrestrial but not an extraterrestrial origin. These observations unambiguously rule out a lunar origin for tektites. Raisbeck also found that other impact-produced materials such as Libyan desert glass contain  $^{26}\text{Al}$  in amounts consistent with what cosmic rays would produce on bombarding terrestrial surface matter for  $\sim 10^7$  years. He suggested, as have others, that the cosmogenic radionuclides may prove useful in the calculation of surface exposure ages of terrestrial rocks.

Tera presented the  $^{10}\text{Be}$  contents of more than 80 australasian tektites and the  $^{26}\text{Al}$  contents of six. Again the  $^{26}\text{Al}/^{10}\text{Be}$  ratios were terrestrial. Tera assessed the small range of values observed for  $^{10}\text{Be}$ , values which average about  $2 \times 10^4$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . Sediments harvested from continental margins exhibit a similar, small range of  $^{10}\text{Be}$  contents. Soils, in contrast, have  $^{10}\text{Be}$  contents that differ markedly from location to location, ranging from  $1 \times 10^4$  to over  $10^6$  atom  $^{10}\text{Be} \text{ g}^{-1}$ . Tera suggests a continental margin for the site of the impact that produced australasian tektites. An origin on a continental margin would accommodate the difficulty specialists have had in identifying a source crater: the crater would have been covered. A margin site may also furnish material



## VGP (cont. from p. 595)

a thesis on sediments is completed also.

Hertzog has pointed out that a tandem Van de Graaf is also operating at the University of Washington (G. Farwell) and that the State University of New York at Stony Brook (A. Chappagne) is starting up. I haven't run across any papers from these facilities, however.

A matter that needs mentioning is the itinerant U.S.-non U.S. research efforts. The first paper reporting AMS radiocarbon dates recently appeared in *Science*. This paper was a collaborative effort by a physicist at San Jose State with the AMS group at the Swiss Federal Institute of Technology. T.-L. Ku and colleagues have been doing beryllium-10 studies of manganese nodules with the Toronto AMS group. It appears to be difficult for all the U.S. scientists who wish to participate in the terrestrial cosmogenic radionuclide work to be involved with U.S. AMS facilities, but at least some of those being left out are being resourceful in finding collaborating facilities somewhere in the world. The problem is not restricted to the U.S. scientists. C. M. Raisbeck of France—the person who started it all and had a 2-year head start—has not been able to get on French facilities, so has been collaborating with researchers at U.S. AMS facilities at the universities of Rochester and, more recently, Pennsylvania. Raisbeck is now getting a General Ionex machine, however.

## Article (cont. from p. 593)

centrifugal accelerations of earth-fixed points, establishment of ties between world datums and determination of polar motion and the earth's gravitational model.

The launching of the last satellite during the IGY and the subsequent launching of dedicated geodetic satellites started a new era of global "datums." DoD was first to solve such a datum in 1960. The solution then became known as the World Geodetic System (WGS) in 1960 thus provided for the first time a "unified" geocentric worldwide coordinate system for global mapping and charting.

## 5. Gravimetric Investigations

Under the IGY gravity program, DoD played two important roles: first, to participate actively in gravity measurements all over the globe and, second, to provide logistic support to other agencies in remote areas like the Arctic and Antarctic.

The measurement program, besides filling in the gaps where gravity observations were scarce and of doubtful accuracy, included the establishment of first-order stations, the calibration of gravimeters, the verification of the connections between established stations, and the extension of the worldwide gravity net including ocean areas. In remote territories like

Elmore has pointed out in me that the very important krypton-81 (not formed by spontaneous fission in nature, like chlorine-36) cannot be done by AMS methods because it does not form negative ions. An alternative, laser-based method is being worked on by a consortium involving Scripps, Bern, and Oak Ridge. G. Wasserburg points out that once improved electron microscopes are developed for the Cameca ion probe, it should be ideal for dating aluminum-26, so AMS might not be necessary for it. Therefore, AMS facilities cannot do all the cosmogenic radionuclides, and the relatively less expensive instruments like the ion probe may be suitable for the time being to do the first order studies in the frontier scientific area of terrestrial cosmogenic radionuclides, their expense (more than \$1000 per day) and lack of high, long-term precision probably will render them unsuitable at some undetermined date in the future. The relatively low General Ionex machine is not yet operating at the 3 MeV mentioned in the specs. Alternatives (desk top cyclotrons of Muller, etc.) are not yet developed. Our problem, therefore, is two fold:

(1) We must get funds for terrestrial cosmogenic radionuclide studies, both for scientific investigations and for operation and development of a few existing AMS facilities. Proposals in scientific exploration (where AMS has the best applications) receive hostile peer review in the earth science area of NSF,

compared to the normal scientific engineering proposals (where you know a lot and can make a good case for the next step).

(2) We must develop less costly, more precise instrumentation over the next decade. As bootlegged research and research funded by other agencies or parts of NSF sponsored research by the Physics Division become published and terrestrial cosmogenic radionuclides begin to move into a less pioneering mode, the hostility in the peer review system of the earth sciences community in NSF may begin to evaporate. This development will stress existing NSF budgets in earth science, especially as the interest of the Physics Division can be expected to wane.

I am a little bit uncertain as to what is the proper thing to do for the next step. Certainly I am not competent to render a Delphic decision. I think two things are needed, however: (1) a strategy on how to proceed over the next 10 years or so in this area and (2) some priorities as to where cosmogenic radionuclides should be, relative to the rest of geochemistry and mineralogy.

As for the importance of this research area, I quote from the previous report of June 16, 1982:

This frontier area of geochemistry is not only exciting—for example, it appears from Be-10 studies that sediments are being subducted in many areas—but has great application to societal goals as well. These applications extend from the dating of mineral deposits (many phosphorite de-

posits thought to be Recent are gradually shown to be Miocene, but they are difficult to date) to the already mentioned ground water dating for waste storage to ground stability evaluation. In ground stability, there is the dating of both sediment and sedimentation rates, evaluation of soils, and dating exposure ages of target rocks. One of the most difficult dating areas involves torrential flooding and land slides and the termination of recurrence intervals. Exposure ages are a good way to get at this if the technique works.

In spite of the difficulties, there are an impressive amount of research, number of papers, and variety of research at the four main facilities (but mainly Rochester and Pennsylvania), as testified to by the symposium organized by G. F. Herzog at the Baltimore AGU meeting on May 30 (EOS, May 3, 1983, p. 282-284), and the meeting report in this issue of *The VGP News*. Papers involving AMS were presented on "the lake systems," "the ice in ice," a beginning on ground water, "the ice in precipitation," "the ice in soil," and also in erosion and deposition, "the ice in phosphorites and in crude petroleum as well as in a manganese crust and, of course, ice in the ice sheets." I've counted correctly, the symposium involved 30 researchers, some of whom are on more than one paper. Their interest is certainly there and should build during the decade.

Bruce R. De  
Editor, The VGP News

Antarctica, gravity measurements were also made in conjunction with seismic and glacial studies to improve our knowledge of isostatic compensation in that area.

DoD was also one of the pioneers that recognized the importance of a "combined" solution from surface gravity information and observations of satellite orbit perturbations available from the IGY earth satellite programs. From the initial computation of the third zonal harmonics in early 1959, DoD was among the first to solve a global geopotential model under the World Geodetic System project the following year.

In its logistics support role, the U.S. Navy's ship *The Compass* (land) provided a gyro-stabilized platform for the first successful sea surface gravity observations on November 29, 1957. This historic operation thus established that it would be possible to acquire data from over 80% of the earth's surface in the succeeding years. Along similar lines, indispensable airborne and ground support were supplied to the IGY Antarctic gravity traverse teams by a U.S. Navy task force.

## 6. Geodetic Laser System

The pioneering efforts to develop satellite laser illuminating and ranging techniques were successfully conducted by AFGL in 1963. Also, experiments with corner-cube retroreflectors (CCR) by the Army Map Service (now

DMA) began in the years following the IGY. These are examples of DoD's indirect contribution to geodesy. It is interesting to note that earlier attempts with CCR's were performed with 60-inch searchlight beams. On the basis of the knowledge acquired as a result of the IGY, it is not a surprise that these attempts failed because of large amounts of backscatter.

As the system design progressed, the CCR's were replaced by more efficient retroreflectors. The original concept to measure distances with geodetic accuracy became a reality. We all know that in today's world lasers form the essential ingredient of some of the most accurate and complex geodetic instruments, and the word laser itself has become synonymous with high accuracy geodesy.

## 7. Summary

On this occasion marking the 25th anniversary of the IGY, this article has recalled some of the DoD activities of the initial years that significantly contributed to the IGY satellite programs involving active, passive, and cooperative satellites.

Today, DoD continues to support the development, enhancement, and application of new technology in such areas as satellite altimetry, airborne gravimetry, gravity gradiometry, inertial surveying, interferometry, charge coupled devices for geodetic astronomy, and others. And, of course, there is a tremendous effort going into the Global Positioning System, which is scheduled to be operational later in this decade.

Time does not allow mention of all the contributions made by DoD components that fulfilled the IGY objectives in satellite geodesy, practical methodology of data reduction, satellite tracking techniques and instrumentation, gravimetric investigations, and other basic research. It suffices to say that much of today's technology is a result of the stimulus provided by the dedicated and concentrated efforts begun during the IGY. It has been a fantastic 25 year era for geodesy and for us.

Contributions require people, and while it is not our intent to mention all of those associated with or supported by DoD during this era, the geodetic world will remain beholden to those represented by the likes of the following: the O'Keefe, Kaula, Chavira, Fischler, Lemys, Anderles, Kersner, Wolf, Golech, Newtons, Cohen, Whipple, Haden, Markowitz, Winkler, Heiskanen, Urdak, Mellers, Rapp, Woodard, Tawani, Williams, Murray, Edwards, Briggs, Whitcomb, Faller, Bender, Thompson, Szabo, Schmidt, Mancini, Gaudin, Martin, Deckers, Baltes, Daugherty, Wilcox, Ewing, Whites, Macomber, Schwiderski.

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## Real-Time Ionosphere Data

Ionospheric electric fields estimated from observations with the STARE (Scandinavian Twin Auroral Radar Experiment) system have been available in real time since August 1983. Such field information is very valuable. To carry out well-defined experiments in the earth's ionosphere/magnetosphere system, it is desirable to have access in real time to data which define the state of this system. On the principle that no geophysical event is unique, conditions in the ionosphere/magnetosphere are recurrent, and thus the state noted by an experiment will recur eventually to determine when experiments have yielded information on currents (magnetometers) and charged particle precipitation (dosimeter, visual auroral data).

STARE consists of two coherent radars with a common field of view covering

200,000 km<sup>2</sup> of the E region over northern Scandinavia. Simultaneous observations of backscatter intensities and Doppler velocities from the two radars allow estimates of the ionospheric electric field with a spatial resolution of 20 x 20 km and a temporal resolution of typically 20 s. The only requirement is that the electric fields exceed a threshold of about 15 mV/m for generation of radar auroral irregularities (see E. Nielsen and J. D. White, *Advances in Space Research*, 2, 131, 1983, and R. A. Greenwald et al., *Radio Science*, 13, 1021, 1978).

The radar stations are computer controlled (Data General Nova 2, 16 Kibibyte memory) and in continuous operation. Both stations are now equipped with communication software (written by C. Steward, Leicester University, UK) and hardware, which permit them to be reached over the normal telephone system. All parameters governing the radar operations as well as all measurements are transmitted at the end of each integration time. This transmission does not interfere with the normal operation of the radar.

All experiments desiring access in real time to STARE data are invited to call one or both of the radar stations and to use the transmitted data to help determine when an experiment should be initiated. Should the STARE system be used later in the analysis and interpretation of observations, then a cooperation with the STARE group should be arranged. We believe that this access could be of interest for ground-based experiments, satellite experiments (for example, HEAT-ING, ELSCAT, VIKING), and rocket and balloon experiments. Two rocket ranges (JAN-DOA and ESRANGE) lie within the STARE field of view.

For more information, contact E. Nielsen, Max-Planck-Institut für Aeronomie, D-3411 Katlenburg-Lindau, FRG. The STARE system is operated by the Max-Planck-Institut für Aeronomie in cooperation with ELAB, University of Trondheim, Norway, and the Finnish Meteorological Institute, Helsinki, Finland.

## Prospecting With Neutrinos

One of the latest attempts to explore the interface between physics and geophysics is the outrageous scheme of Alberto De Rújula, Sheldon Glashow, Robert Wilson, and Georges Charpak, to be published in *Physics Reports*. In what these theoretical and experimental physicists described recently as "our mad project" (*Physics Today*, August 1983), a high-energy neutrino beam is to be used as a geophysical prospecting tool.

The beam would be able to look for oil, natural gas, and high-atomic-number metals, and it would be able to probe the vertical density distribution of the earth. De Rújula et al. come to this project from the world of big physics machines, so it is natural to expect that the "Geotom," the field instrument to supply and focus the neutrino beam, is to be big also.

To their credit, the four particle physicists have worried about the consequences and the cost. They think that the machine could be built for about \$1 billion, or even less; however, they concede their real concern is that it would be even more: "It might cost very much more if built under the besting supervision of present-day bureaucrats. The construction could well be drawn out into a job that will run for years as the rust, doubt, and then double again. This sobering ideal for projects is not, experience informs us, the exception."

What the neutrino beam really does is to provide a train of oscillating disturbances and nuclear particles along its path as it travels through the earth. The process functions as follows: "When a nucleus hits a nucleus it produces a forward moving shower of charged particles (which ionize the medium)

and neutral particles; the neutrals (mainly pions) decay to produce additional ionizing radiation. This sudden deposition of energy, produced in a narrow cone of ionization, produces an acoustic signal."

The neutrinos, whose energies are measured in TeV, are produced by a Geotron, which is an underwater-based synchrotron. The Geotron would provide a proton beam of 10-20 TeV that in turn would impinge upon a target to yield a collimated beam of pions and kaons that would undergo decay as mesons to neutrinos. This process of conversion would be tightly collimated in a tunnel called a Snout. The Snout is actually a flexible tube that holds a series of superconducting magnets. Even at a distance of 1000 km the low divergence of the neutrino beam would result in a radius of only 10 m.

GENIUS (Geological Exploration by Neutrino-Induced Sound), is the application of the Geotron to locate petroleum deposits at great distances. At 1000 km an array of geophones with frequency ranges of 1-100 Hz would be tuned in on the acoustic beam and at the same time filter background and seismic noise. Enormous areas can be surveyed and at great depth (the signal energy falls off with the first rather than the fourth power of depth, as with seismic waves).

It also would be feasible, according to the proposed plan, to measure the muon beam. This GENIUS would become GEMINI (Geological Exploration with Muons Induced by Neutrino Interaction). Muon sensors could detect differences as the beam passed through high-atomic-number metal concentrations.

A crucial property of the beam in all applications is that it can penetrate the entire earth. In the Geotom a special, vertically oriented beam of neutrinos would be aimed at the earth in such a way that it would travel through the center of the earth, penetrating the core. Muon detectors would sample the beam at various angles. It should be possible to obtain a determination of the earth's radial density distribution to high accuracy in a short time.

The project is being proposed as a selling point for construction of the Texatron, the huge hadron collider accelerator that has been suggested for the Texas A&M campus. The Texatron could provide a neutrino beam that would be used to test these ideas of geophysical exploration.—PMB

## China's Petroleum Reserves

Perhaps because of declining yields inland, the People's Republic of China has moved to its source-rich coast to develop additional petroleum reserves. During 1979 and 1980, 44 foreign oil-exploration companies engaged in what has been termed the world's record "seismic shoot" over 11,000 km<sup>2</sup> extending from the Yellow Sea through the South China Sea, and including Hainan Island (Gulf of Tonkin), the bay lying east of Hanoi and west of Hainan Island. These offshore oil reserves are estimated to be 40-100 billion barrels.

The seas off the mainland are relatively shallow (most drilling has been done in less than 100 m) but they are straggly. According to a recent description, "Typhoon can . . . occur in the area at almost any time of the year, and the strong winds (160 km/h or more) which they generate frequently wreak havoc in the Philippines, Vietnam, China, and occasionally, Hong Kong. Typhoon Vera, which hit southern China in late July, did immense damage and claimed dozens of lives. The main implications for oil operation, naturally relate to rig design and safety measures but onshore facilities will also have to be designed accordingly" (*New Scientist*, Sept. 8, 1983).

Exploration and drilling techniques similar to those used in the North Sea, although at shallower depths, are being used. The politi-

cal boundaries will not be so easily settled, however. The China Sea and its islands are a great potential source of political conflict, the entire area being claimed by mainland China, the Philippines, Taiwan, Vietnam, and others.—PMB

## Society Merger Rejected

Members of the American Society of Photogrammetry (ASP) and members of the American Congress on Surveying and Mapping (ACSM) have rejected a plan to consolidate the two societies.

An affirmative vote of at least two-thirds of each society's members present or represented by proxy was required for the merger. Voting in favor of the consolidation plan were 61.2% of ASP voters and 63.7% of ACSM voters. Almost 50% of eligible members of each society participated in the September 21 vote.

The consolidation had been discussed for several years and was formally proposed in June 1981. In September 1981 the two societies agreed to affiliate (EOS, November 10, 1981, p. 765) and agreed to prepare a consolidation plan. Although the consolidation was rejected last month, the societies will retain their affiliation. As affiliated societies, they are located in the same building in Falls Church, Va., and share a single governmental affairs program and an educational program.

## JOI Seeks Nominations

Joint Oceanographic Institutions (JOI), Inc., is seeking nominations for a new panel to coordinate scientific ocean drilling. To be called the U.S. Science Advisory Committee (USAC), the panel would complement the work of the Joint Oceanographic Institutions for Deep Earth Sampling (JOIDES).

The Board of Governors of JOI says it wants USAC to include broad representation from academic, government, and industry sectors outside the 10 JOI member institutions, and is inviting nominations, including a brief vita and list of publications, for membership on USAC to be sent, by November 3, to John H. Garwin, JOI, Inc., 2100 Pennsylvania Ave., N.W., Suite 316, Washington, DC 20037.

## Geophysicists

Richard C. Hart is the new secretary of the U.S. National Committee for the Committee on Space Research (COSPAR). Hart succeeds Dean Kastel, who is now the executive secretary of the National Research Council's Space Science Board. In addition to his new duty with COSPAR, Hart serves as the staff officer for the Space Science Board committees on space astronomy and astrophysics, solar and space physics, and data management and computation.

John H. McElroy was recently appointed assistant administrator for the National Environmental Satellite, Data, and Information Services of the National Oceanic and Atmospheric Administration. McElroy had been acting assistant administrator since February 1983.

Timothy D. Steele has accepted a position as water resources manager of the Denver office of In-Situ, Inc., a high-technology consulting company serving the mining and energy industries. Steele previously worked for more than 13 years with the U.S. Geological Survey's water resources division. Most recently, he was chief of the water quality group in the Denver office of Woodward-Clyde Consultants.

## Recent Ph.D.'s

See periodically this information on recently accepted doctoral dissertations in the disciplines of geophysics. Faculty members are invited to submit the following information on individual listed, above the signature of the faculty advisor or department chairman: the dissertation title, author's name, name of the degree-granting department and institution, and month and year degree was awarded. If possible include the current address and telephone number of the degree recipient (this information will not be published).

*Method and Radiative Isotopes in Submarine Hydrothermal Systems*, Kyung-Ryul Kim, Scripps Institution of Oceanography, Univ. of California, San Diego, June 1983.

*The Oceanographic and Geologic Components of San Surface Topography*, Victor Zlotnicki, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, June 1983.

*Velocity and Upwelling Near an Isolated Feature on the Continental Shelf*, Stephen M. Chiswell, Marine Sciences Research Center, State Univ. of New York, Stony Brook, August 1983.

## Climatic Changes

by M.I. Budyko (1977)  
English translator, R. Zolna  
English translation editor, L. Levin

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*Effects of Sea, Age and Photoperiod on Hypomorphology in Brook Trout *Salvelinus fontinalis**, Stephen D. McCormick, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, September 1983.

*Formulation of Time Elements: Uptake, Isotopism, and 100 m.y. Paleoclimatic History*, Margaret L. Delaney, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, September 1983.

*Processing and Interpretation of Arctic Ocean Refraction Data*, Gregory L. Duckworth, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, September 1983.

*Purification and Characterization of the Hepatic Microsomal Aroclor 1248 System for the Coastal Atlantic Fish *Stenotomus chrysops**, Alan V. Klotz, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, September 1983.

*Wave-Induced Turbulent Flow Near a Rough Bed: Implications of the Time-Varying Eddy Viscosity*, John H. Trowbridge, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, September 1983.

*Effects of Pore Pressure, Confining Pressure and Partial Saturation on Permeability of Sandstones*, Joel Dan Walls, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., January 1983.

*Relative Motion Between Oceanic and Continental Plates in the Pacific Basin*, David Cal Engebretson, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., January 1983.

*Velocity, Attenuation, and Natural Fractures in Shallow Basaltic*, Daniel B. Moos, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., January 1983.

*A Study of the Seismic Structure of Upper Oceanic Crust Using Wide-Angle Reflections*, Kristin M. Roler, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, February 1983.

*Spectral Reflectance of Near-Earth Asteroids: Implications for Composition, Origin and Evolution*, Lucy A. McFadden, Dept. of Geology and Geophysics, Hawaii Institute of Geophysics, Univ. of Hawaii, May 1983.

*Dip-Awareness by Fourier Transform*, Ira David Hisk, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., June 1983.

*Flow and Skin Friction over Natural Rough Beds*, Christopher Paola, Joint Program in Oceanography and Oceanographic Engineering, WHOI/MIT, June 1983.

*Migration of Reflection Seismic Data in Angle-Midpoint Coordinates*, Richard Albert Olden, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., June 1983.

*New Techniques in the Analysis of Geophysical Data Modelled as a Multichannel Autoregressive Random Process*, P. A. Tyrakis, Laboratory in Applied Geophysics, Dept. of Mining and Metallurgical Engineering, McGill Univ., Montreal, Canada, June 1983.

*Spatial Coherence of Seismic Waveforms*, Keith McLaughlin, Dept. of Geology and Geophysics, Univ. of California, Berkeley, June 1983.

*Wave Propagation in Porous Rock and Models for Crustal Structure*, Perry Dean Jones, Dept. of Geophysics, School of Earth Sciences, Stanford Univ., June 1983.

## News

## Antarctic Research Priorities Set

Highest priority for future research under the U.S. Antarctic Program should be given to the "extraction of the unique climatic record preserved in the Antarctic ice sheet" and to "the study of the response of marine life to the unique environment at the edge of sea ice." That's the word from the National Research Council's (NRC) Polar Research Board, which was asked by the National Science Foundation (NSF) to recommend priorities for research efforts in the Antarctic. NSF funds the U.S. Antarctic Program. The board ranked eight large-scale research projects in order of priority, identified three smaller, more specific projects that should be included in the U.S. Antarctic Program, and listed other necessary, supportive activities with "wide-spread implications and applications."

The board says that the best long-term results will be gleaned from the U.S. Antarctic Program with a mixture of the "intensive, integrated, large-scale projects focused on one or more of the principal research questions," the "smaller-scale, lower-cost projects," and "other activities, including ongoing collection and analysis of data, publication of scientific results, and production of maps."

In addition to those eight projects, the board recommended three small projects. The first, designated as very high priority, is studying the life-history patterns and adaptations of the Antarctic biota. The remaining two, assigned high priority, are measuring the heat budget at and around South Pole Station and studying the biogeochemical processes in Antarctic ecosystems.

Other activities within the purview of the U.S. Antarctic Program that the Polar Research Board said should be continued include "the production of topographic and geologic maps; the collection of meteorites; and the monitoring programs at South Pole and McMurdo stations that provide data on upper atmospheric, cosmic, and solar phenomena, earthquakes, and earth tides, constituents of the atmosphere, and standard meteorological data."

To effectively implement the research projects identified, the NRC board said that "support systems permitting longer operating seasons and covering wider geographic areas" are required.

Charles R. Bentley, at the University of Wisconsin's Geophysical and Polar Research

standing of global energy transfer in the magnetospheric cusp and polar cusp.

• An interdisciplinary investigation of the structure and intensity of the Weddell gyre and the impact of the associated fluxes on the climatic, glacial, and biological environment.

High priority was assigned to

• A coordinated program of geophysical studies aimed at understanding West Antarctic crustal structure and history and the dynamics of the ice sheet.

• A multidisciplinary study integrating physical and biological measures to determine the causes for, and ecological consequences of, the swarming behavior of krill.

• A study of the interaction between electromagnetic waves and energetic particles, making use of conjugate-point measurements.

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Charles R. Bentley, at the University of Wisconsin's Geophysical and Polar Research

Center, is chairman of the Polar Research Board. Three ex officio members and four agency liaison representatives complement the 16 scientists who constitute the board. W. Timothy Hushen is executive secretary.

## Diamonds at High Pressure

New calculations indicate that diamonds may not have any phase changes at pressures below  $23 \times 10^{11}$  N m<sup>-2</sup> (23 Mbar) (M. T. Yin and M. I. Cohen, *Physical Review Letters*, 50, 2006, 1983). The importance of this result of course affects the applications of diamonds beyond their value as gem stones. Single crystal diamonds have served as very strong windows for the transmission of a range of radiation spectra in high-pressure experiments. The ultimate strength of diamonds, however, could mark the limit of experiments at 1 Mbar and higher pressures.

A few laboratories have been able to experiment at pressures equivalent to those at the earth's mantle-core boundary (approximately 1.5 Mbar), but such experiments can hardly be considered routine. Extrapolations from experiments done by the General Electric Research Laboratory in Schenectady, N.Y., and by J. A. Van Vechten (*Physica Scripta*, 7, 261, 1971) suggested that diamonds might undergo a phase change at about 1.5 Mbar; if so, diamonds used as windows would fall at the transition. There has been speculation that the phase change would include a reordering of diamond's electronic structure to the metallic state.

Yin and Cohen calculated that a high-pressure transition to the simple cubic structure would occur in diamonds at 23 Mbar. The pseudopotential technique they employed has proven successful in numerous other cases and is becoming a valuable tool in crystal cal-

culations. According to an account by J. Willis of the Churche Laboratory, University of Oxford, "Yin and Cohen assumed only the atomic number of the element and the type of crystal structure and then determined the lattice constants, cohesive energies and bulk moduli" (*Nature*, 303, 102, 1983).

The earlier work of Van Vechten had pointed toward a B4-in (tetragonal) crystal structure for the high-pressure carbon phase, in analogy with silicon and germanium. Yin and Cohen obtained consistent results but a different result for the diamond. The 23 Mbar transition pressure borders well for Mbar maximum. The new results say nothing, however, of the yield strength of diamond, which could result in flow or failure at considerably lower pressures.—PMB







**Geophysical Tenure-Track Appointment/Department of Geology, University of Toledo.** The position is effective September 1, 1984. Individuals with strong backgrounds in exploration geophysics—applied geophysics are of primary interest although other specializations will be considered. The Ph.D. is required as well as a strong commitment to effective teaching and research. The department has modern facilities and offers M.S., B.A., and M.S. degrees to approximately 60 undergraduate and 30 graduate students. The faculty consists of eight full time and five adjunct professors actively involved in a wide range of research pursuits. Interested persons should submit a letter of application, resume, transcripts, and three letters of recommendation to: Sam L. Dean, Chairman of Search Committee, Department of Geology, University of Toledo, Toledo, Ohio 43606, phone (419) 537-2240 or (419) 527-2009. University of Toledo is an equal opportunity/affirmative action employer.

**Professor of Marine Geophysics/Tenure-Track/Stanford University.** The Department of Geophysics is seeking candidates for a tenure track position in the broad area of marine geophysics and tectonics. We seek a creative scientist with experience in geology, tectonics, and synthesizing marine geophysical data and whose research interests cover deep-sea, igneous, and tectonic processes on oceanic plates and continental margins. Inquiries are invited from marine geophysicists with demonstrated scientific record in one of the above subjects of marine geophysics or tectonics, who have demonstrated an ability to develop new ideas and research directions, and to guide and teach graduate and undergraduate students. In considering this appointment we are interested in maximizing interactions with ongoing research groups in marine geology, plate tectonics, paleomagnetism, tectonics and regional geology at Stanford. One new faculty member will be expected to develop a strong research program involving both government and industrial participation.

Salary and rank will be commensurate with experience and background. Please submit a resume, a brief description of teaching and research interests, and references to:

Dr. Anna Nui  
Department of Geophysics  
321 Mitchell Building  
Stanford University  
Stanford, CA 94305

Stanford University is an equal opportunity employer, and encourages the application of qualified women and minorities.

## POSITIONS WANTED

**Physical Chemistry.** Ph.D. specialized in isotopic Geochemical Basic Research would consider challenging opportunity. P.O. Box 018 American Geophysical Union, 2000 Florida Avenue, N.W., Washington, DC 20009.

## STUDENT OPPORTUNITIES

## GRADUATE STUDENT

**NASA TRAINEESHIPS**  
The Florida State University is accepting applications from prospective graduate students for participation in its NASA sponsored Traineeship Program in Oceanographic Remote Sensing Techniques and Physics of Air-Sea Interaction. The stipend for the calendar year is \$10,000. Students may be enrolled for a degree in either oceanography or meteorology. For further information or application, please write:

Dr. James J. O'Brien  
NASA Traineeship Program  
Meteorology Annex  
The Florida State University  
Tallahassee, Florida 32306  
(904) 644-4581

## Meetings

## Announcements

## River Basins

The International Water Resources Association (IWRA) will hold a seminar at Linköping University in Sweden June 4-8, 1984, to discuss the relevance of the river basin approach to land and water management. The seminar will try to reach conclusions about plans for future actions using the river basin as the basic unit, specifically with regard to the criteria for environmental planning, for conflict resolution, and for developing coordinated land and water control.

Papers will deal with eight selected river basins and with four issues: river basins as ecosystems; legal and administrative issues; the problems of growing urban systems; and the problems of coordinating control and management of land and water resources.

Those interested in attending should contact Ulrik Lohm, Water Theme, Linköping University, S-58183, Linköping, Sweden.

## Offshore Minerals

A symposium to discuss plans for assessing and developing mineral resources in the recently proclaimed Exclusive Economic Zone off the coast of the United States will be held November 15-17 at the U.S. Geological Survey (USGS) National Center in Reston, Va. The symposium is being sponsored by the USGS, the Minerals Management Service, and the Bureau of Mines to aid in organizing a coordinated government, university, and industry effort to evaluate the potential mineral resources in the new zone; the zone was proclaimed on March 10, 1983, to extend U.S. mineral rights 200 miles offshore.

Presentations on current federal marine mineral resource activities will be followed by panel discussions on the science of resource assessment, the engineering technology involved, and the legal ramifications in developing these offshore minerals. The Secretary of the Interior is scheduled to be the keynote speaker; others from the White House, Congress, Department of Commerce, Navy, industry, and academia also will speak.

For additional information, contact the Assistant Secretary, Energy and Minerals, Department of the Interior, Washington, DC 20240 (telephone: 202-343-5691).

## AGU Fall Meeting: Travel, Housing, Registration

The 1983 Fall Meeting of the American Geophysical Union will be held in San Francisco, Calif., December 5-9, at the Cathedral Hill Hotel and the Holiday Inn Golden Gateway Hotel. San Francisco is a dynamic, exciting city, known to the world for its spectacular scenery, fabulous restaurants, cosmopolitan life style, and gentle climate. It is a superb meeting location at any time of the year.

## Registration

Everyone who attends the meeting must register. Preregistration (received by November 10) saves you time and money. The fee will be refunded to you if AGU receives written notice of cancellation by November 28. Registration rates are as follows:

	Preregistration	After Nov. 10
Member	\$65	\$80
Student member	\$32	\$47
Retired senior member (65 or over)	\$32	\$47
Nonmember	\$90	\$105
Student nonmember	\$41.50	\$56.50

Registration for 1 day only is available at one half the above rates, either in advance or at the meeting. Members of the American Congress on Surveying and Mapping, the American Meteorological Society, the American Society of Photogrammetry, the European Geophysical Union, and the Union Geofisica Mexicana, may register at the AGU member rates.

If you register as a nonmember for more than 1 day, the first-year dues for joining AGU will be waived if a completed application is received at AGU by Feb. 10, 1984. To preregister, fill out the registration

## Hotel Accommodations

Blocks of rooms (\$47 singles, \$53 doubles) are being held at the Cathedral Hill, the Holiday Inn Golden Gateway, the Holiday Inn Civic Center, the San Francisco, and the Grosvenor Inn for those attending. Read the housing application, and mail the completed application form to the housing bureau early to ensure reservations at your preferred hotel. *Reservation forms must be sent directly to the Housing Coordinator, AGU Fall Meeting, San Francisco Housing Bureau, P.O. Box 5612, San Francisco, CA 94101.* Do not send housing reservation forms to the hotel. Reservations must be received by November 10.

## RETURN THIS FORM WITH PAYMENT TO:

Meeting Registration  
American Geophysical Union  
2000 Florida Avenue, N.W.  
Washington, DC 20009

PLEASE PRINT CLEARLY

NAME ON BADGE

AFFILIATION

MAILING ADDRESS

TELEPHONE #

HOTEL

Days you plan to attend

Please check the appropriate box(es)

☐ Dec. 5 ☐ Dec. 6 ☐ Dec. 7  
☐ Dec. 8 ☐ Dec. 9

Please check appropriate box

☐ Member AGU ☐ Nonmember

Members of the cooperating societies may register at AGU member rates

Member cooperating society:

☐ AMS-American Meteorological Society  
☐ ASP-American Society of Photogrammetry  
☐ ACSM-American Congress on Surveying and Mapping  
☐ EGU-European Geophysical Union  
☐ UGM-Union Geofisica Mexicana

Nonmembers

If you register as a nonmember for more than 1 day, the first-year dues for joining AGU will be waived if a completed application is received at AGU by Feb. 10, 1984.

Preregistrants

Your receipt will be in your preregistration packet. The registration fee will be refunded if written notice of cancellation is received in the AGU office by November 28. The program and meeting abstracts will appear in the November 8 issue of *Eos*.

AGU 1983 FALL MEETING  
DECEMBER 5-9  
San Francisco, California

## REGISTRATION FORM

Deadline for Receipt of  
Preregistration  
NOVEMBER 10, 1983

(rates applicable only if received by November 10 with payment)

	More than one day	One day
MEMBER	<input type="checkbox"/> \$65 <input type="checkbox"/> \$80	<input type="checkbox"/> \$32 <input type="checkbox"/> \$47
STUDENT MEMBER	<input type="checkbox"/> \$32 <input type="checkbox"/> \$47	<input type="checkbox"/> \$16 <input type="checkbox"/> \$21
RETIRED SENIOR MEMBER*	<input type="checkbox"/> \$32 <input type="checkbox"/> \$47	<input type="checkbox"/> \$16 <input type="checkbox"/> \$21
NONMEMBER	<input type="checkbox"/> \$90 <input type="checkbox"/> \$105	<input type="checkbox"/> \$45 <input type="checkbox"/> \$52.50
STUDENT NONMEMBER	<input type="checkbox"/> \$41.50 <input type="checkbox"/> \$56.50	<input type="checkbox"/> \$20.75 <input type="checkbox"/> \$28.25

\*65 or over

## SECTION LUNCHEONS/DINNER

Circle section and indicate number of tickets. All lunches begin at noon. SPR dinner begins at 6:30 P.M.

- ☐ Planetology/Volcanology, Geochemistry and Petrology, Tuesday, \$9  
☐ Seismology/Tectonophysics, Tuesday, \$5  
☐ Ocean Sciences and Paleomagnetism, Wednesday, \$5  
☐ Hydrology, Wednesday, \$9  
☐ Ocean Sciences, Wednesday, \$9  
☐ Solar-Planetary Relationships, Wednesday, \$20 (dinner)  
☐ Atmospheric Sciences, Thursday, \$9  
☐ Geodesy, Thursday, \$9

Total Enclosed \$

(All orders must be accompanied by payment or credit card information. Make check payable to AGU.)

Charge to: ☐ American Express ☐ Visa ☐ MasterCard

Card Number

Master Card Interbank No.

Expiration Date

Signature

Office Use

Clerk	_____
Check	_____

## FIELD TRIP FORM

I wish to attend the Franciscan Nummulite field trip on Sunday, December 4. My check for \$25 is enclosed.

In case I am not among the first 40:

☐ I wish to be put on the waiting list. (If you don't go, money will be returned on the day of the trip.)

☐ I wish my money returned.

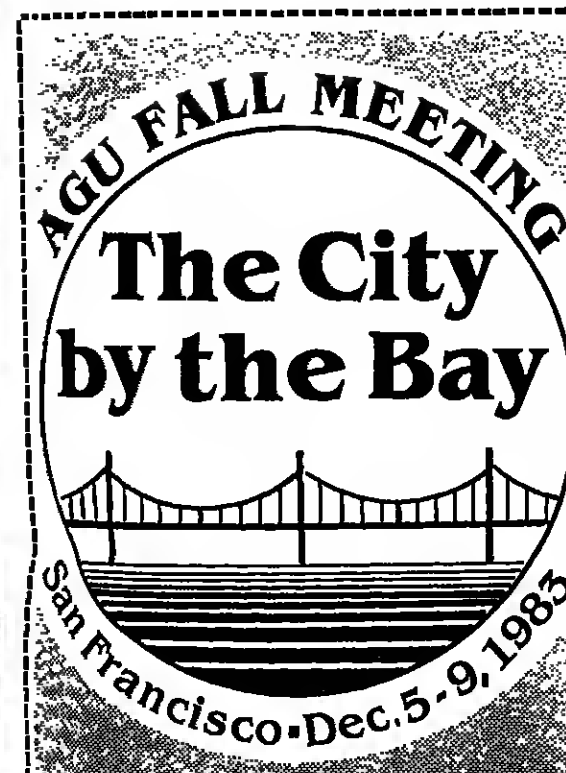
Signature \_\_\_\_\_ Print Name \_\_\_\_\_

Date

Address

Telephone

Mail form to: M. C. Blinke, Jr., Mail Stop 75, U.S. Geological Survey, 345 Middlefield Road, Menlo Park, CA 94025



## HOTEL ACCOMMODATIONS

## PARTICIPATING HOTELS

Cathedral Hill Hotel  
Van Ness at Geary Street  
(800) 227-4730

Holiday Inn Golden Gateway  
1500 Van Ness Avenue  
(415) 441-4000

Grosvenor Inn  
Van Ness and Geary  
(415) 673-7411

Holiday Inn Civic Center  
50 8th Street  
(415) 626-6103

San Francisco Hotel  
1231 Market Street  
(415) 626-8000

## ROOM RATES FOR ALL HOTELS

Single \$47  
Double \$53  
Twin \$53

Suites available upon request

**PARKING:** Cathedral Hill Hotel: free to registered guest  
Holiday Inn Golden Gateway: free to registered guest  
San Francisco Hotel: free to registered guest

All hotel reservations must be made on the housing form by November 1, 1983. No telephone requests will be accepted. Confirmations will be mailed directly to registrants by the individual hotels. After confirmation has been received, changes and cancellations should be made directly to the hotel.

Mail your completed form directly to:

Housing Coordinator  
AGU Fall Meeting  
San Francisco Housing Bureau  
P.O. Box 5612  
San Francisco, CA 94101

ber 1 to be confirmed. Do not write or call AGU for room reservations.

Free parking is available only to registered guests of each hotel as indicated.

## Scientific Sessions

The program summary appears later in this issue. The preliminary program and the abstracts will be published in *Eos*, November 8. The final program, with presentation times, will be distributed at the meeting. Both the Cathedral Hill and the Holiday Inn Golden Gateway hotels will be used for all disciplines.

**Poster Sessions** Poster sessions will be held throughout the meeting in the Eldorado Room at the Cathedral Hill Hotel. Check the program for detailed scheduling. AGU will provide each poster-session presenter with a mounting area measuring 4 x 6 feet (1.25 x 2 m). Plan your exhibit to fit this space. The board will be assigned by number corresponding to the presenter's abstract number. The boards will be set up in the Eldorado Room before the poster session begins. Thumb racks, push pins, tape, and scissors will be available in the meeting room.

## Exhibits

The exhibits will be located on the Mezzanine, Cathedral Hill Hotel, Monday, December 5, through Thursday, December 8, 9:30 A.M. to 4:00 P.M.

The following exhibitors are confirmed:

Academic Press, Inc.  
American Geophysical Union  
Defense Mapping Agency/HTC  
Digital Imaging Processing  
EG&G Geometrics  
Elsevier Science Publishing Co.  
Hammar  
Jet Propulsion Laboratory  
Kinematics  
Nature's Own  
Phoenix Geophysics  
Qualimetrix, Inc.—WEATHERtronics  
Refraction Technology  
Sciencemeter Instruments Co.  
Springer-Verlag, New York  
Tetra Tech  
Terra Technology  
U.S. Geological Survey

## Social Events

An icebreaker party on Monday evening on the Mezzanine at the Cathedral Hill Hotel will be the opening social event of the meeting. To honor John W. Handin, the 1983 recipient of the Bocher Medal, and those 1983 AGU Fellows who were not present at the 1983 Spring Meeting, there will be an awards ceremony and wine tasting reception on Thursday evening, 6:00-7:30 P.M., in the

Meetings (cont. on p. 602)

American Geophysical Union  
1983 FALL MEETING

## HOUSING REGISTRATION FORM

READ CAREFULLY and RETURN FORM DIRECTLY TO THE SAN FRANCISCO HOUSING BUREAU AT THE FOLLOWING ADDRESS:

Housing Coordinator  
AGU Fall Meeting  
SF Housing Bureau  
P.O. Box 5612  
San Francisco, CA 94101

Please print or type all information, abbreviating as necessary. Confirmation will be sent by the hotel to the individual named in Part I. If more than one room is required, this form may be photocopied.

## Part I

REQUESTOR

Last Name \_\_\_\_\_ First \_\_\_\_\_

Name of Company or Firm \_\_\_\_\_

Street Address or P.O. Box Number \_\_\_\_\_

City \_\_\_\_\_ State/Prov. \_\_\_\_\_ Zip-U.S.A. \_\_\_\_\_

Country \_\_\_\_\_ Telephone Number \_\_\_\_\_

## Part II

INSTRUCTIONS: Select **THREE** hotels of your choice from the list of participating facilities, then enter the name on the lines below.

First Choice

Second Choice

Third Choice

NOTE: Rooms are assigned on a "First Come, First Served" order, and if none of your choices is available, another facility will be assigned based on a referral system. A cut-off date is in effect; your application may not be processed if received after 14 days prior to your arrival date. AGU housing registration deadline is November 1.

## Part III

INSTRUCTIONS: 1. Select type of room desired with arrival and departure dates.  
2. **PRINT** or **TYPE** names of **ALL** persons occupying room.  
3. If more than two persons share a room, check twin and the hotel will assign two double beds.

CHECK ONE		Guest Names (Last name first)	
<input type="checkbox"/> SINGLE (Room with one bed one person)	Arrival Date _____	1. _____	_____
<input type="checkbox"/> DOUBLE (Room with one bed two persons)	Arrival Time _____ AM/PM	2. _____	_____
<input type="checkbox"/> TWIN (Room with two beds two persons)	Departure Time _____	3. _____	_____
<input type="checkbox"/> EXTRA PERSON		4. _____	_____

IMPORTANT NOTE: Hotel MAY require a deposit or some other form of guaranteed arrival. If so, instructions will be on your confirmation form.











